

# *Energy Calibration and Stability of the ANKA Storage Ring*

A.-S. Müller, I. Birkel, E. Huttel, F. Pérez\*, M. Pont\*, R. Rossmanith

Institut für Synchrotronstrahlung – ISS  
Forschungszentrum Karlsruhe

\* now at ALBA Synchrotron Light Source, Spain

## *Overview*

- Brief introduction to ANKA
- Why energy stability is an issue at ANKA
- How to calibrate the beam energy in a very precise (but simple) way
  - resonant depolarisation
  - makeshift polarimeter
- Summary



## Research & development in various fields:

- environmental analysis
- medicine/bio technology
- material science
- microsystem technology
- (astro) particle physics
- IT science
- nano technology
- ...



# Ångströmquelle Karlsruhe

- Tradition in the production of electromagnetic radiation in KA



1887

First observation of electromagnetic waves in Karlsruhe by Heinrich Hertz  
("Hertz Dipole")

1996

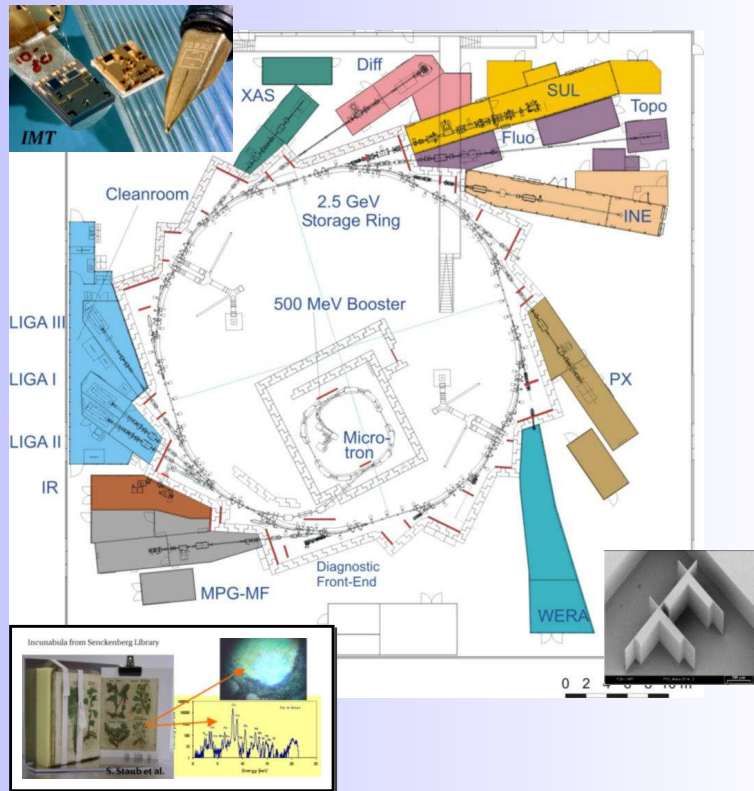
Ground-breaking ceremony for ANKA  
First beam in 2000  
Regular user operation since 2001/2002



# The ANKA Facility



- Electron storage ring with a beam energy of 2.5 GeV
- hosts 12 beamlines
  - running: 6 + 3
  - commissioning: 3
  - planned: 1
- Mostly bending magnet beam lines
  - free space for 5 IDs
  - two installed
- Development of SC IDs
  - planar SCU to come in the near future
  - first prototype of helical SCU



10/12/2004

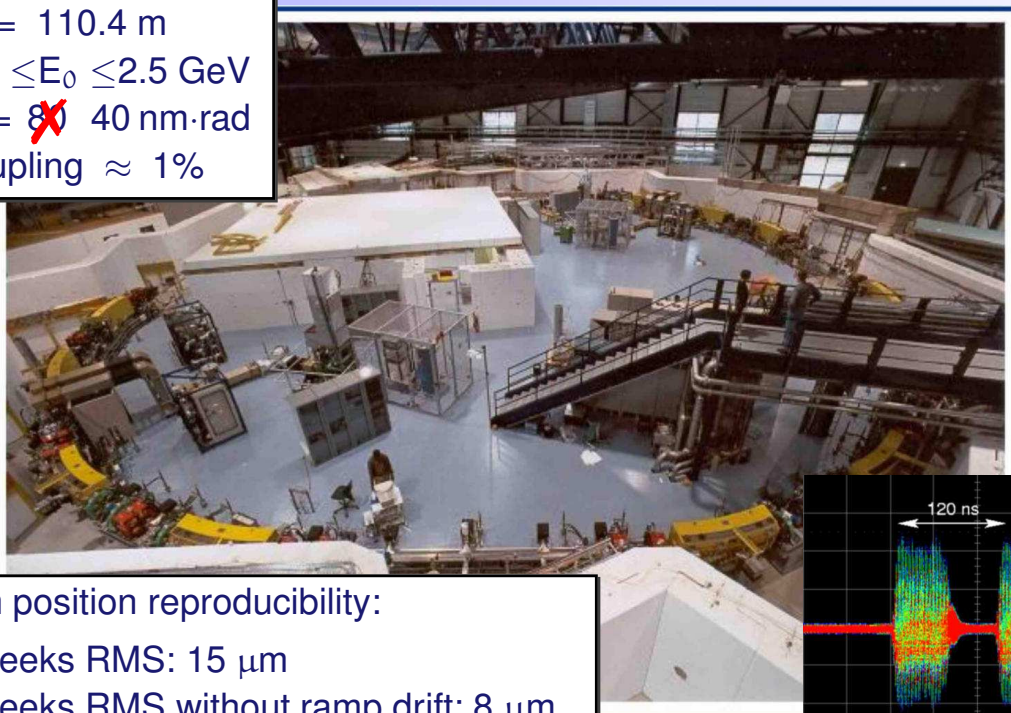
Energy Calibration of ANKA – Introduction

IWS2004 – page 5/21

# The ANKA Storage Ring

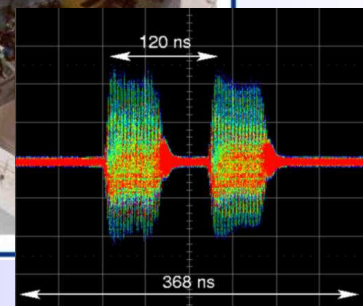


- ✓ C = 110.4 m
- ✓  $0.5 \leq E_0 \leq 2.5$  GeV
- ✓  $\epsilon_x = 8 \times 40$  nm·rad
- ✓ coupling  $\approx 1\%$



Beam position reproducibility:

- ✗ 6 weeks RMS: 15  $\mu\text{m}$
- ✗ 6 weeks RMS without ramp drift: 8  $\mu\text{m}$
- ✗ 1 h RMS: < 2  $\mu\text{m}$



10/12/2004

Energy Calibration of ANKA – Introduction

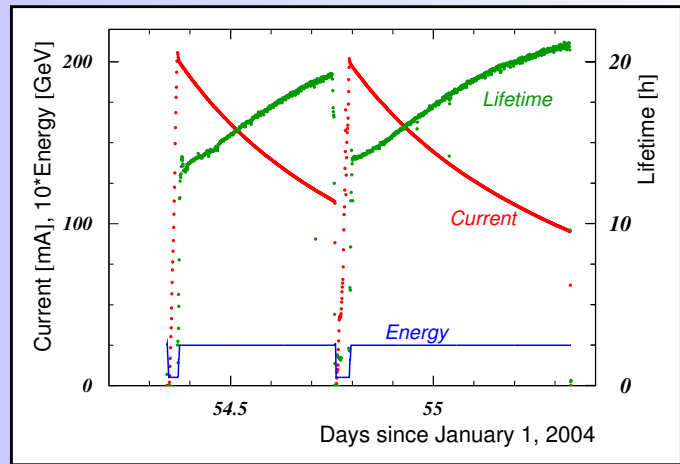
IWS2004 – page 6/21



# A Typical Day in ANKA's Life



- A typical day at ANKA sees two injections of  $\approx 200$  mA: re-fills in the morning and late afternoon
- Regular user operation at 2.5 GeV
- Special operation mode: beam at 1.3 GeV (e.g. for lithography)
- This means:
  - large current changes
  - cycling of bends and quadrupoles
  - ramp ( $B_{\text{dipole}}$  range from 0.3 to 1.5 T)
- Compensate thermally induced drifts in the bending field ( $E_0$ )



10/12/2004

Energy Calibration of ANKA – Introduction

IWS2004 – page 7/21

# Field Drift Correction



- 30 beam position monitors
- 28 horizontal correctors
- 16 vertical correctors
- SVD based orbit correction by cosylab
- choice between measured or model response matrix
- RF frequency as a free parameter in the correction

$$\frac{\Delta p}{p} = - \frac{1}{\alpha_c} \frac{(f_{RF} - f_{RF}^c)}{f_{RF}}$$

HORIZONTAL CORRECTION				VERTICAL CORRECTION					
AVG	RMS	STD	MAX	AVG	RMS	STD	MAX		
last true	-0.263	0.716	0.866	1.422	last true	-0.119	0.553	0.540	1.375
averaged	-0.011	0.191	0.190	0.445	averaged	-0.002	0.053	0.124	
saved av.	0.000	0.000	0.000	0.000	saved av.	0.000	0.000	0.000	

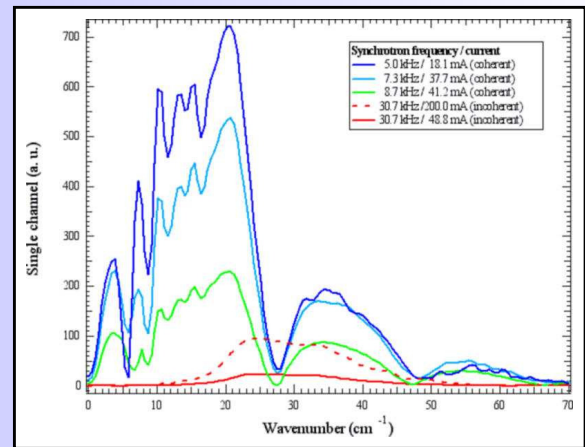
s [m]	Corrector	angle	correct	after	s [m]	BPM	pos	avg pos	saved p.	delta avg
1.6	MCH_S1_01	0.18709	0.09551	0.27727	3.5	DBPM_S1_01	0.054	0.053		
6.7	MCH_S1_02	0.01080	0.23445	0.24317	6.4	DBPM_S1_02	-0.122	-0.119		
9.5	MCH_S1_03	-0.01253	0.22817	0.21653	8.8	DBPM_S1_03	-0.053	-0.052		
13.3	MCH_S1_04	-0.07105	0.29428	0.22431	11.5	DBPM_S1_04	0.447	0.445		
18.9	MCH_S1_06	0.17218	0.21596	0.38958	15.6	DBPM_S1_05	-0.238	-0.228		
20.7	MCH_S1_07	0.13903	0.13019	0.27150	18.5	DBPM_S1_06	-0.215	-0.216		
25.0	MCH_S1_08	-0.06351	-0.14741	-0.20844	20.9	DBPM_S1_07	0.345	0.345		
29.3	MCH_S2_01	0.32269	0.15099	0.47267	23.6	DBPM_S1_08	-0.162	-0.162		
34.3	MCH_S2_02	0.18284	-0.08813	0.09503	31.1	DBPM_S2_01	0.097	0.096		
36.1	MCH_S2_03	0.02487	-0.18096	-0.15473	34.0	DBPM_S2_02	-0.252	-0.251		
40.4	MCH_S2_04	0.32529	0.01613	0.34323	36.4	DBPM_S2_03	0.166	0.166		
					43.2	DBPM_S2_05	-0.047	-0.046		
					46.1	DBPM_S2_06	0.026	0.026		
					48.5	DBPM_S2_07	-0.077	-0.071		

10/12/2004

Energy Calibration of ANKA – Introduction

IWS2004 – page 8/21

- Since spring 2004 new operation mode with reduced  $\alpha_c$  for short bunches
  - production of CSER in the FIR range
  - intensity enhancement of up to 5 orders of magnitude
  
- small changes in corrector strengths lead to visible energy changes
  - orbit correction with RF frequency as free parameter necessary for reproducible conditions



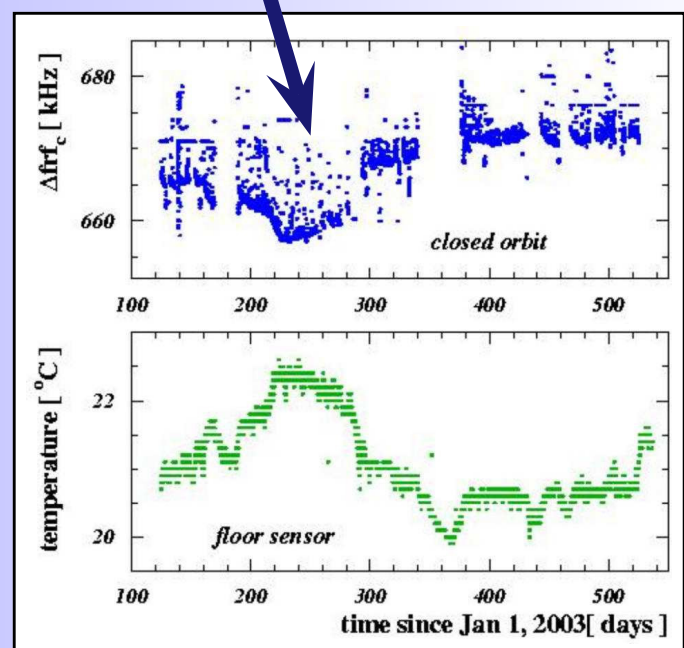
Y.-L. Mathis et al.

# External Influence

- RF frequency adjustment keeps path length constant

$$\frac{\Delta f_{RF}}{f_{RF}} \propto \frac{\Delta C}{C} \propto - \frac{\Delta E}{E}$$

- daily drifts caused by injection and ramping scheme
- significant seasonal dependence
- how well does this truly reproduce the nominal beam energy?



- Why do we need to know the beam energy?
  - as a cross-check for the energy correction by frequency adjustment
  - for ID characterisation:  
 $\epsilon_\gamma \propto E^2 \Rightarrow 1\% \Delta E/E \rightarrow 2\% \Delta \epsilon_\gamma / \epsilon_\gamma$
  - for further development of optics, an excellent knowledge and understanding of the present optics is needed: energy errors are gradient errors!
  
- Measure  $E_0$  with highest possible precision: **Resonant Depolarisation**



- Polarisation build-up by emission of synchrotron radiation:
  - asymmetry in spin flip probability leads to transverse polarisation
  - max. polarisation is given by size of asymmetry term:

$$8/5\sqrt{3} \approx 92.4\%$$

- Polarisation level increases exponentially with build up time

$$\tau_p = \frac{8}{5\sqrt{3}} \frac{m_0^6 c^{10} \rho^3}{\hbar r_c E_0^5}$$

- for ANKA at 2.5 GeV:  $\tau_p \approx 10$  minutes

- The motion of the spin vector  $\vec{s}$  of a relativistic electron in the presence of electric and magnetic fields  $\vec{E}$  and  $\vec{B}$  is described by the Thomas-BMT (Bargmann, Michel, Telegdi) equation:

$$\frac{d\vec{s}}{dt} = \vec{\Omega}_{\text{BMT}} \times \vec{s}$$

- The spin precession frequency  $\vec{\Omega}_{\text{BMT}}$  can be written as

$$\vec{\Omega}_{\text{BMT}} = -\frac{e}{\gamma m_0} \left[ (1 + \alpha\gamma) \vec{B}_{\perp} + (1 + \alpha) \vec{B}_{\parallel} - \left( \alpha\gamma + \frac{\gamma}{1 + \gamma} \right) \vec{\beta} \times \frac{\vec{E}}{c} \right]$$

- $\alpha = (g_e - 2)/2 = 0.001159652193(10)$
- The average over all particles of the number of spin oscillations per revolution is defined as the spin tune

$$\nu = f_{\text{spin}}/f_{\text{rev}} = \frac{(g_e - 2)/2}{m_0 c^2} E_0$$

# Resonant Depolarisation

- Horizontal magn. field  $B_x$  modulated with  $f_B$  is applied to the beam
- For a certain phase relation between the kicks of the depolariser and the spin tune the small spin rotations add up coherently from turn to turn and the polarisation is destroyed

→ resonance condition for spin rotations

$$f_{\text{dep}} = (k \pm [\nu]) \cdot f_{\text{rev}}$$

- To determine the spin tune, the frequency of the depolariser field is slowly varied with time over a given frequency range

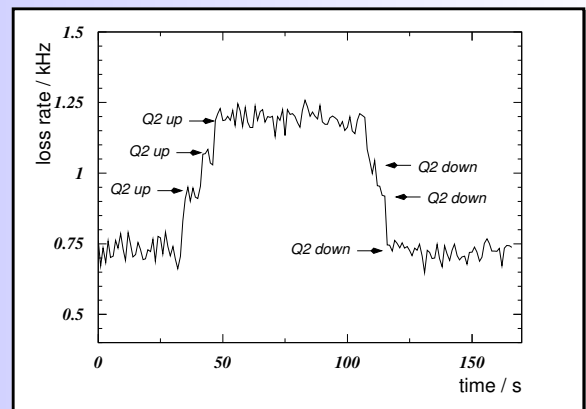
→ precision:  $\Delta E/E \approx 10^{-5}$

# Polarisation and Loss Rate



- Touschek cross-section depends on electron beam polarisation
  - use particle loss rate as a measure for polarisation level
- Pioneering work by
  - P. Kuske et al., "High Precision Determination of the Energy at BESSY II", EPAC 2000
  - C. Steier et al., "Energy Calibration of the Electron Beam of the ALS Using Resonant Depolarisation", EPAC 2000
  - S.C. Leeman et al., "Precise Beam Energy Calibration at the SLS Storage Ring", EPAC 2002
- Set up a beam loss monitor in a Touschek sensitive region...

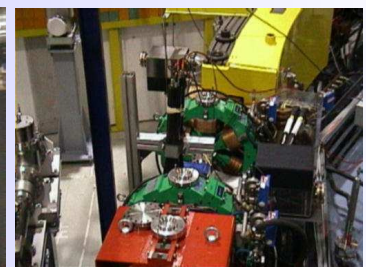
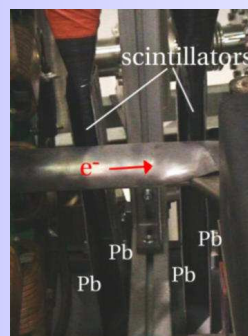
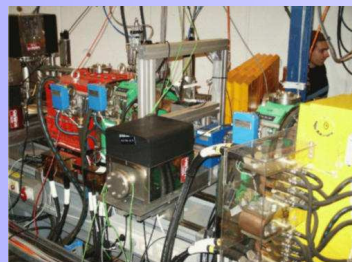
(test of detector sensitivity) →



# Polarisation Detection

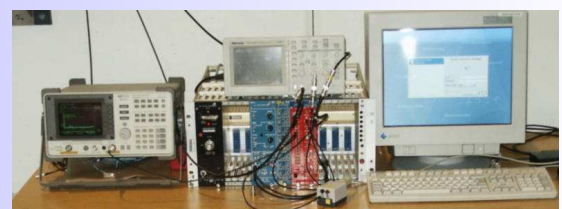


- First attempts with scintillators wrapped in lead sheets to suppress the contribution of synchrotron radiation to the count rate



- Final setup: Pb-Glass block with photo multiplier

- Photo multiplier pulses are converted to NIM signals and counted using a custom made interface to a Linux PC





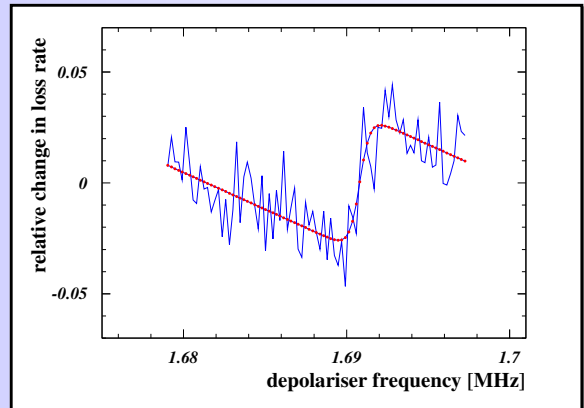
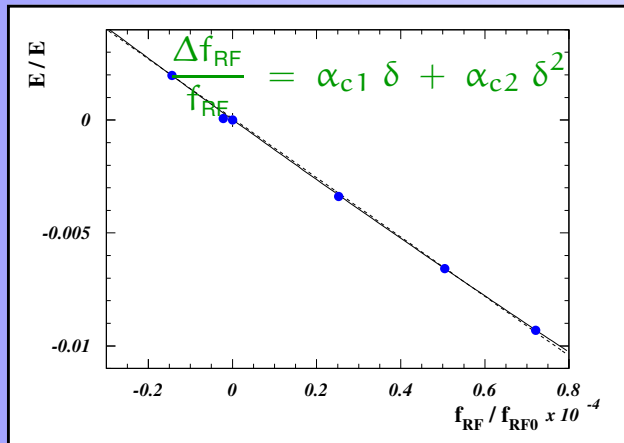
# Some RDP Results



- Fit to loss rate scans using

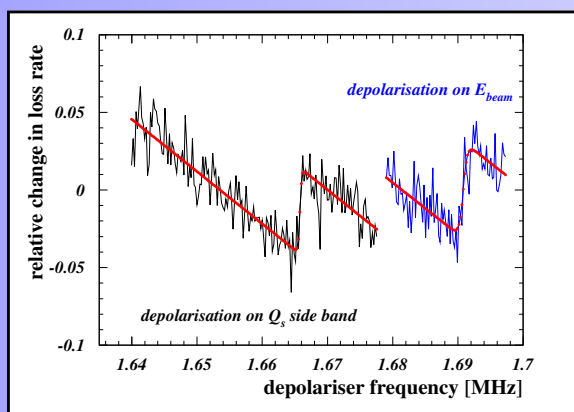
$$r = a - \frac{\partial r_I}{\partial t} t + \frac{\Delta r}{1 + \exp\left\{-\frac{t-t_d}{\sigma_d}\right\}}$$

- Measure  $\alpha_c$  using RDP scans for different  $f_{RF}$ :



- Expectation for linear term:  $7.2 \cdot 10^{-3}$
- $\alpha_{c1} = (7.39 \pm 0.01) \cdot 10^{-3}$
- Quadratic term:  $\alpha_{c2} = (4.1 \pm 0.2) \cdot 10^{-2}$
- highly significant but small

# Bunch Lengthening & RDP

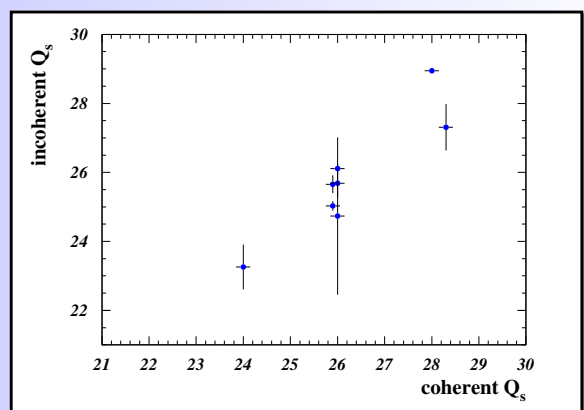


- Depolarisation can also occur on synchrotron side bands
- Due to the single particle nature of the depolarisation process, this happens at  $f_{dep}/f_{rev} = [\nu] \pm Q_s^{inc}$
- RMS bunch length:  $\sigma_s \propto 1/Q_s^{inc}$

- The coh./inc ratio is a measure for bunch lengthening with current

$$\frac{Q_s^{coh}}{Q_s^{inc}} = 1 - \lambda I_{bunch}$$

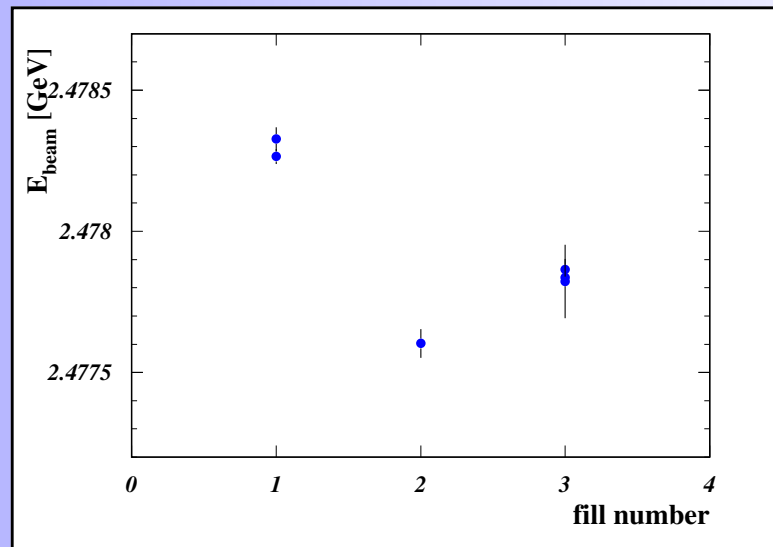
- Average coherent/incoherent ratio:  $(1.02 \pm 0.03)$
- no significant lengthening



# Energy Reproducibility



- Fill-to-fill variations of the order of  $10^{-4}$
- Reproducibility within single fill  $\approx 10^{-5}$



- Energy drift compensation by RF frequency adjustment seems to work reasonably well

# Summary



- The beam energy is an important parameter for accelerator and insertion devices ✓
- Very precise measurement possible with resonant depolarisation ✓
- Detection of polarisation change with Touschek-loss-polarimeter ✓
- Energy stability acceptable:  $\mathcal{O}(10^{-4})$  ✓

