



A users viewpoint: absorption spectroscopy at a synchrotron

Frithjof Nolting

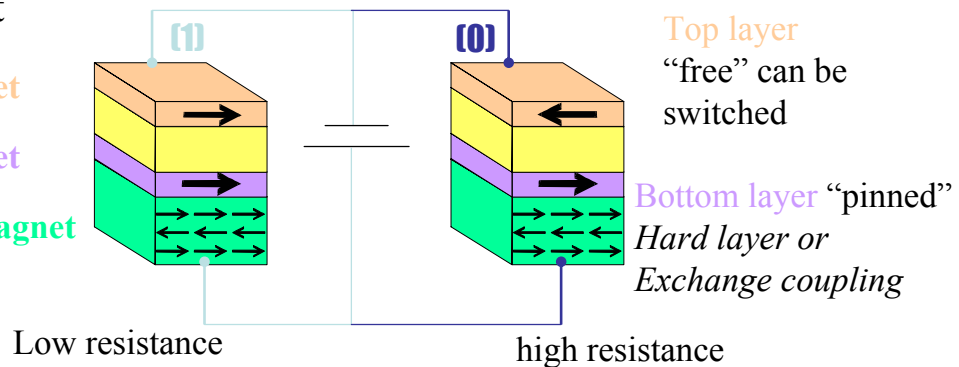
Magnetic data storage and recording

GMR Element

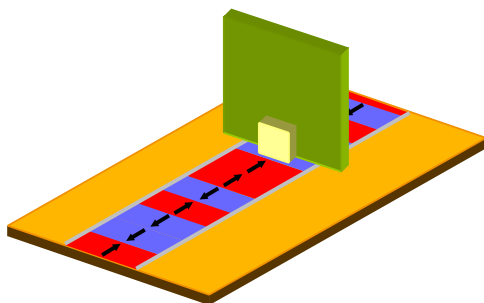
Ferromagnet

Ferromagnet

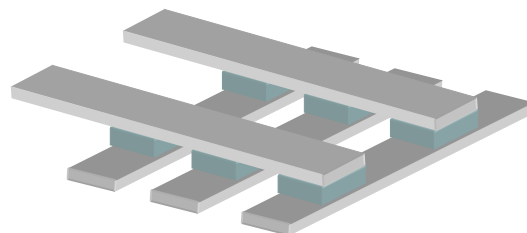
Antiferromagnet



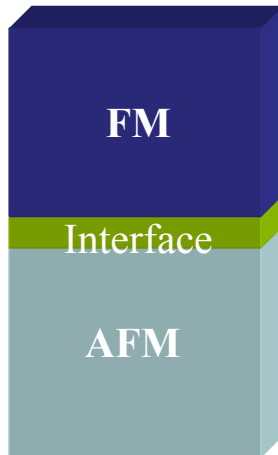
Hard disk head



MRAM



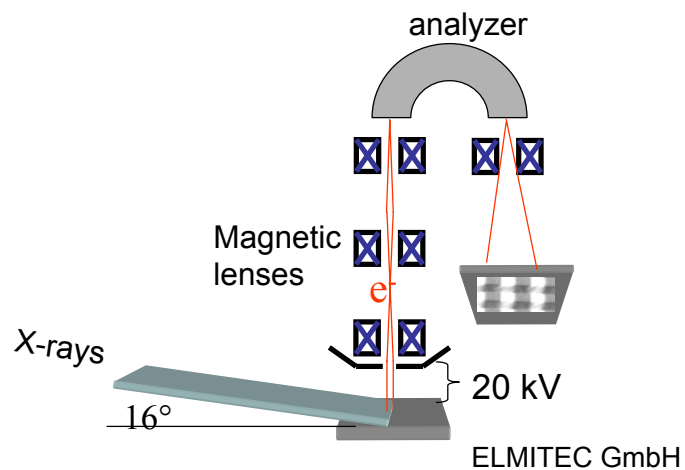
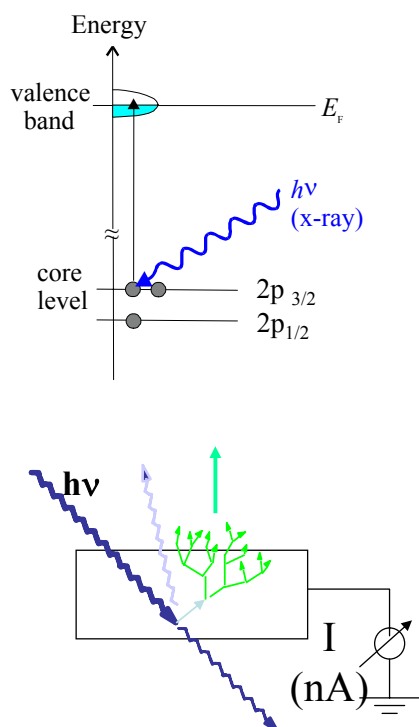
- Different models for AFM/FM coupling exist.
- Different assumptions on AFM structure lead to complete different results.
- Spin arrangement at the interface is important



- High Spatial Resolution
- Elemental Sensitivity
- Ferromagnetic Contrast
- Antiferromagnetic Contrast
- Surface/Interface Sensitivity

X-ray absorption spectroscopy (XAS) with spatial resolution
Photoemission Electron Microscope (PEEM)

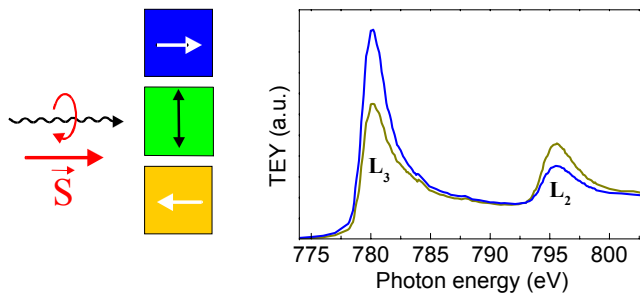
Photoemission Electron Microscope



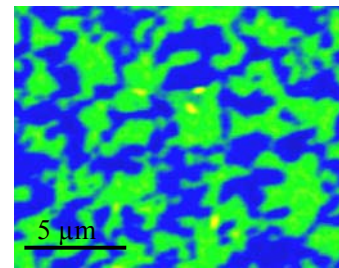
Sensitive to:

- elemental composition
- chemistry
- structural parameters
- electronic structure
- magnetic properties

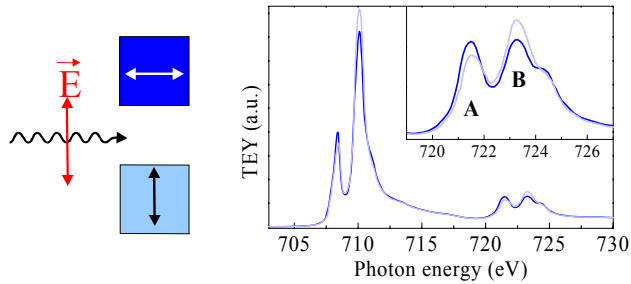
XMCD (X-ray Magnetic Circular Dichroism)



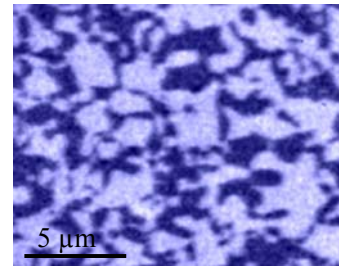
Co



XMLD (X-ray Magnetic Linear Dichroism)



LaFeO₃

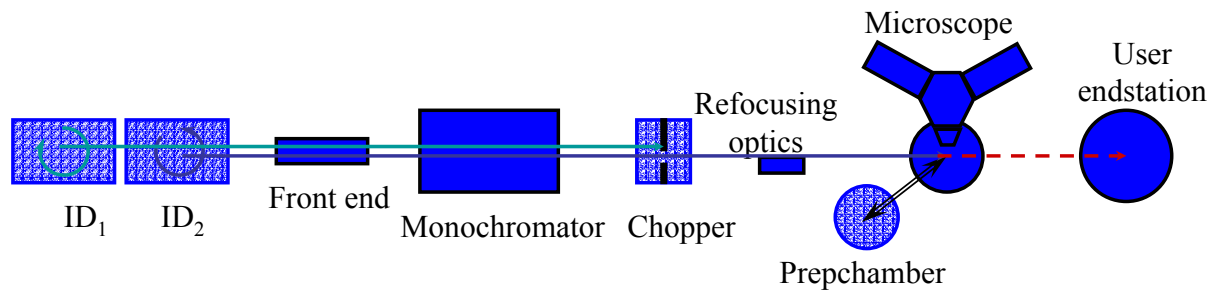


J. Stöhr *et al.* Science 1993

A. Scholl *et al.* Science 2000

F. Nolting *et al.* Nature 2000

SIM Beamline Layout



Undulator

T. Schmidt

- **2 Elliptical undulators**
- Pure permanent magnet
- **95eV < hv < 2000eV**
- $>10^{19}$ photons/s/mrad²/mm²/400mA
- 100 % circular polarization [125 - 900 eV] reduced on higher harmonics
- Hor. & vert. linear polar.

Optics

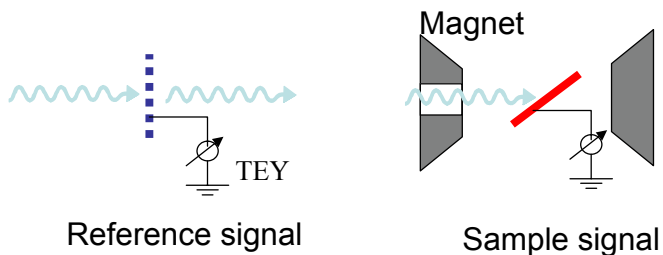
U. Flechsig

- Plane grating monochromator
- $E/\Delta E > 8000$
- $<5\%$ 2nd order light
- **Switch helicity**
- Focus $30 \times 100 \mu\text{m}^2$

Endstation

C. Quitmann & F. Nolting

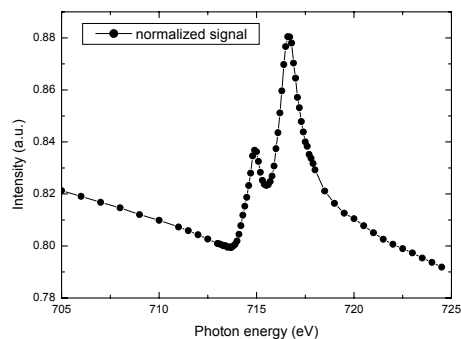
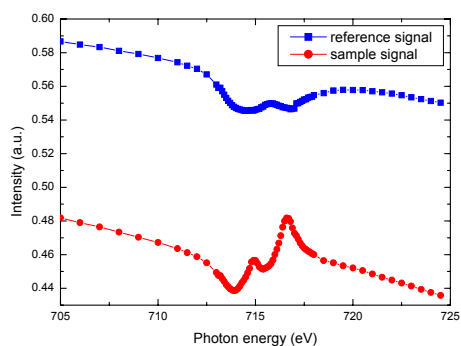
- SLS endstation:
 - **PEEM & LEEM**
 - $\Delta x \sim 25 - 50 \text{nm}$ spatial
 - $\Delta E \sim 150 \text{meV}$ energy
 - Sample Prep chamber
 - **User endstation**



Sample

= Norm

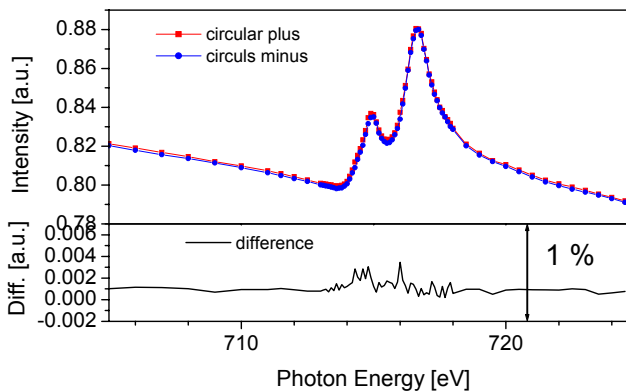
Reference



Moving gap and monochromator, stop, measure (1s – 1minute), moving ...

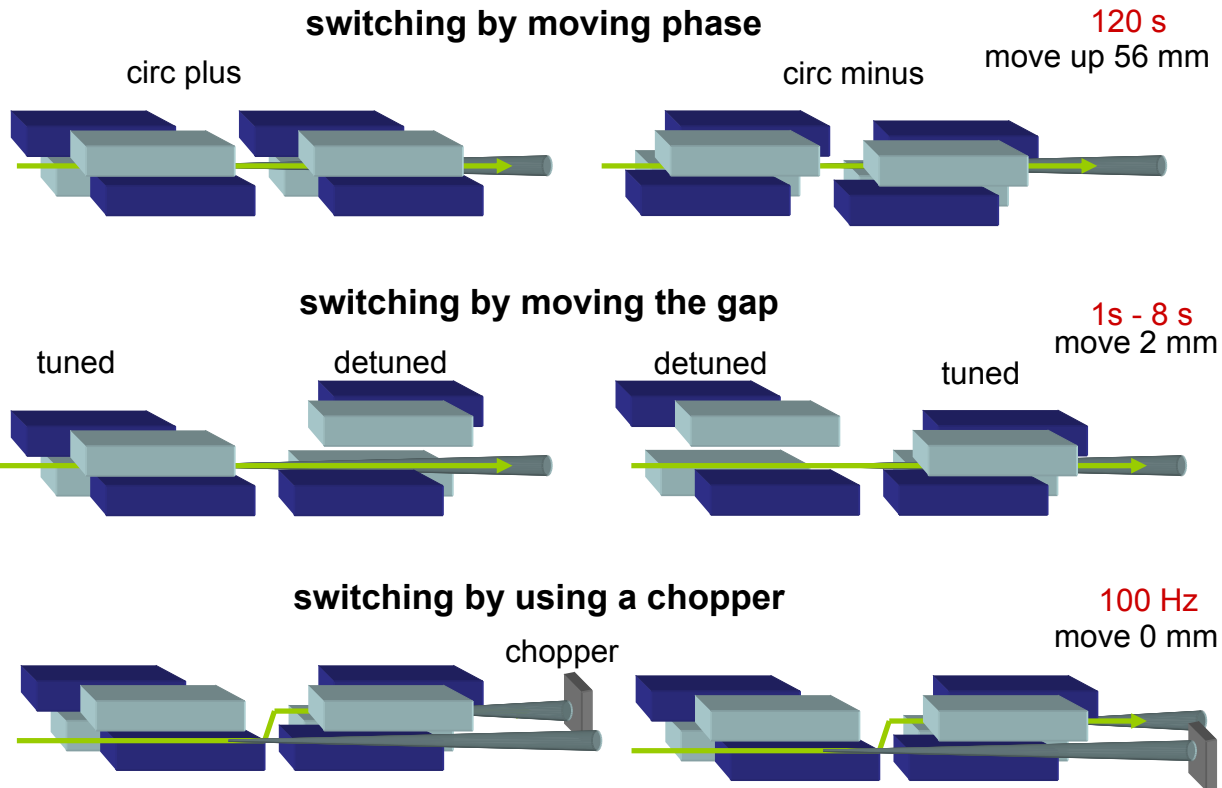
How do we measure

- Change polarization and repeat
- Take difference of spectra



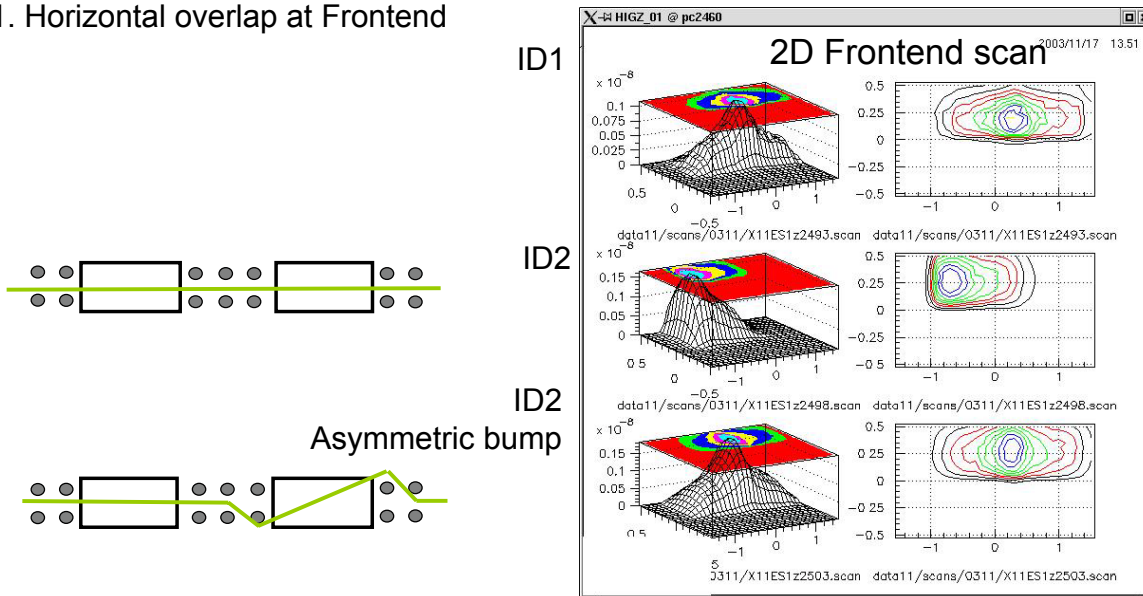
Absorption spectrum requires frequent moving of gap and shift

**must not effect other beamlines
transparent!**

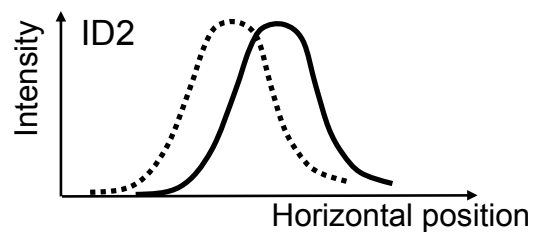
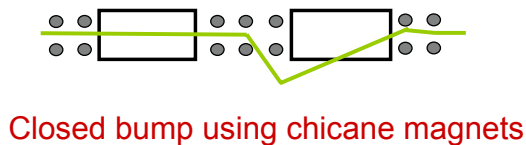


Alignment of IDs - Horizontal

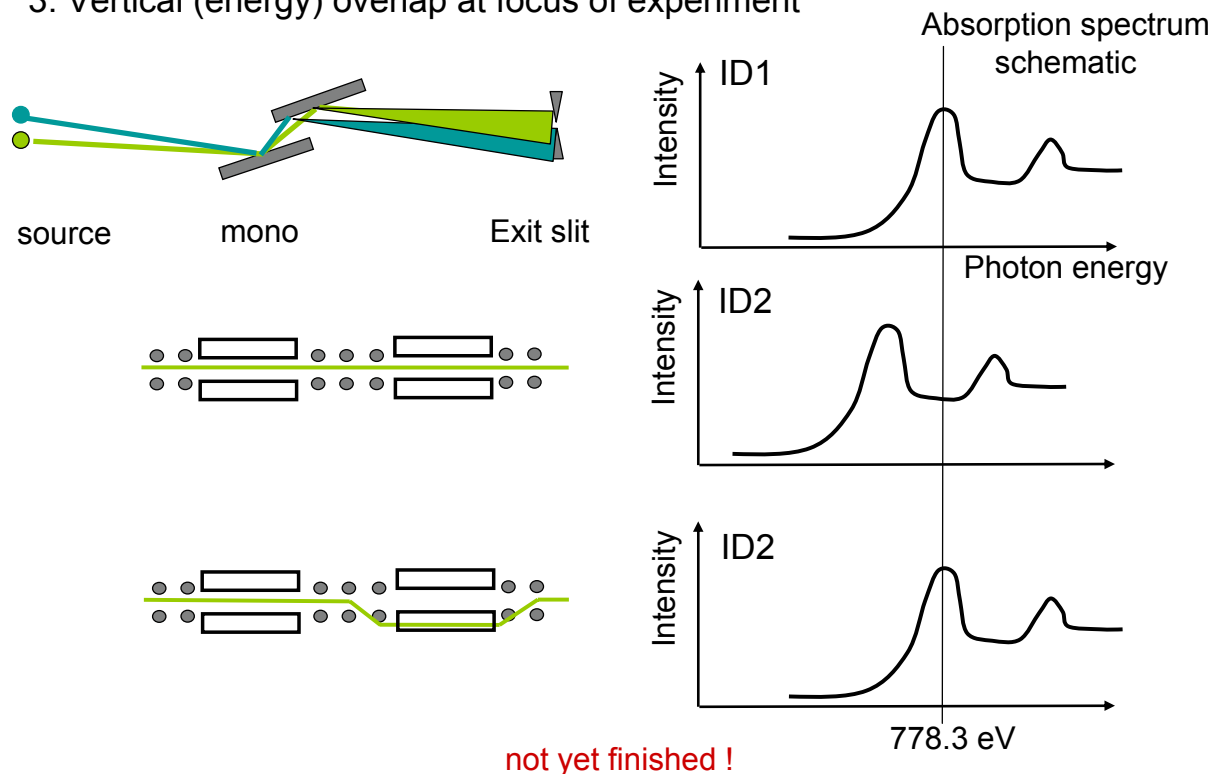
1. Horizontal overlap at Frontend



2. Horizontal overlap at focus of experiment



3. Vertical (energy) overlap at focus of experiment



Beam variation - Noise

horizontal movement \longrightarrow intensity variation

XAS normalization reduces it by a factor of 10-100

PEEM no normalization

days	no problem
hours	no problem
sec - minutes	bad
mili seconds	ok

10 μ m about 2%

vertical movement \longrightarrow energy variation

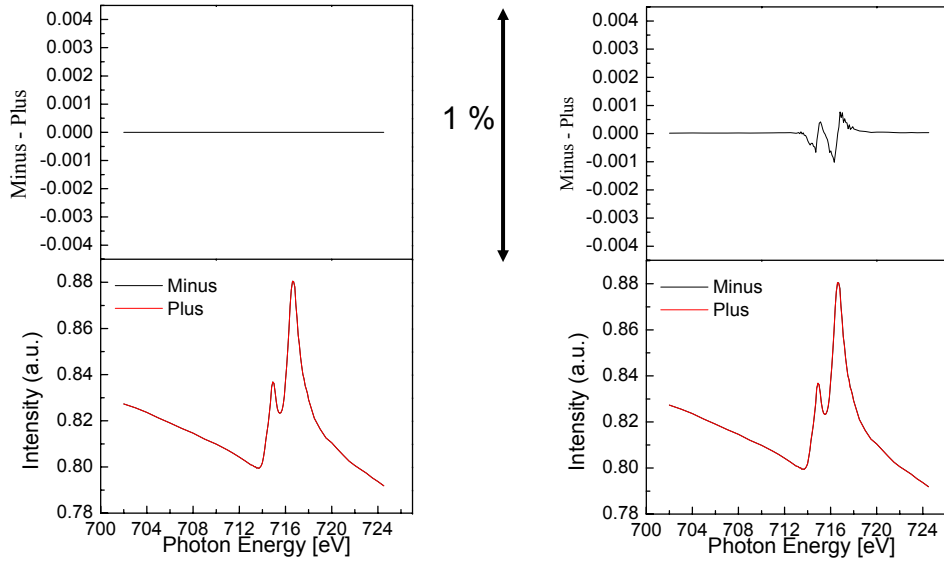
no normalization!

days	no problem
hours	bad
seconds	bad
mili seconds	ok

10 μ m about 10 meV

Identical spectra

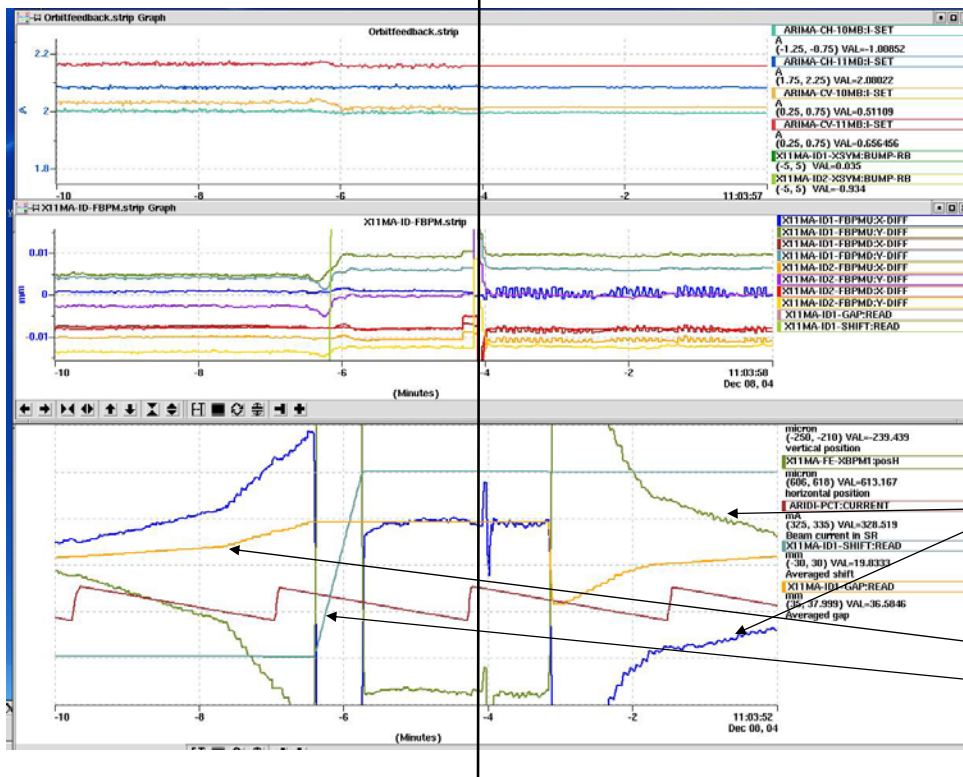
Energy shift of 10 meV



Orbit feedback

Fast orbit feedback

slow orbit feedback



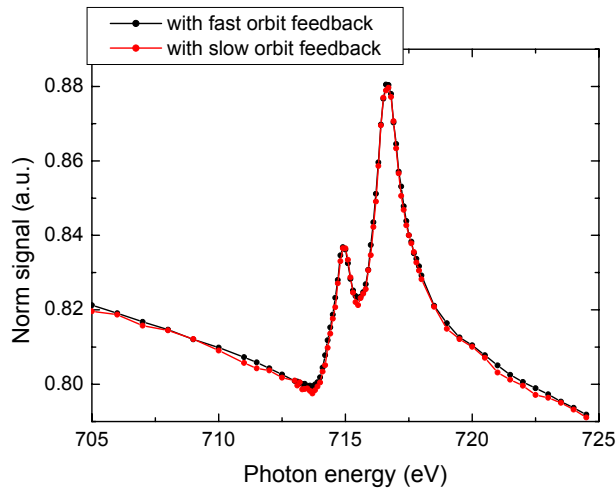
Orbit correctors?

beam position
in ID
(Bergoz)

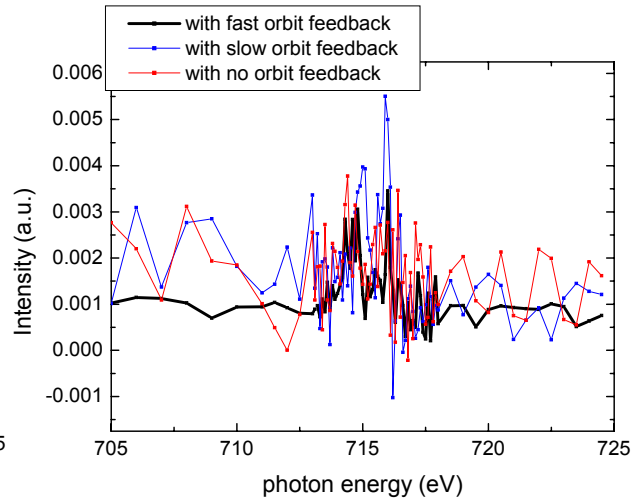
X-ray position
in beamline

Shift and gap

Normalized signal, circular plus



Difference circular plus and minus

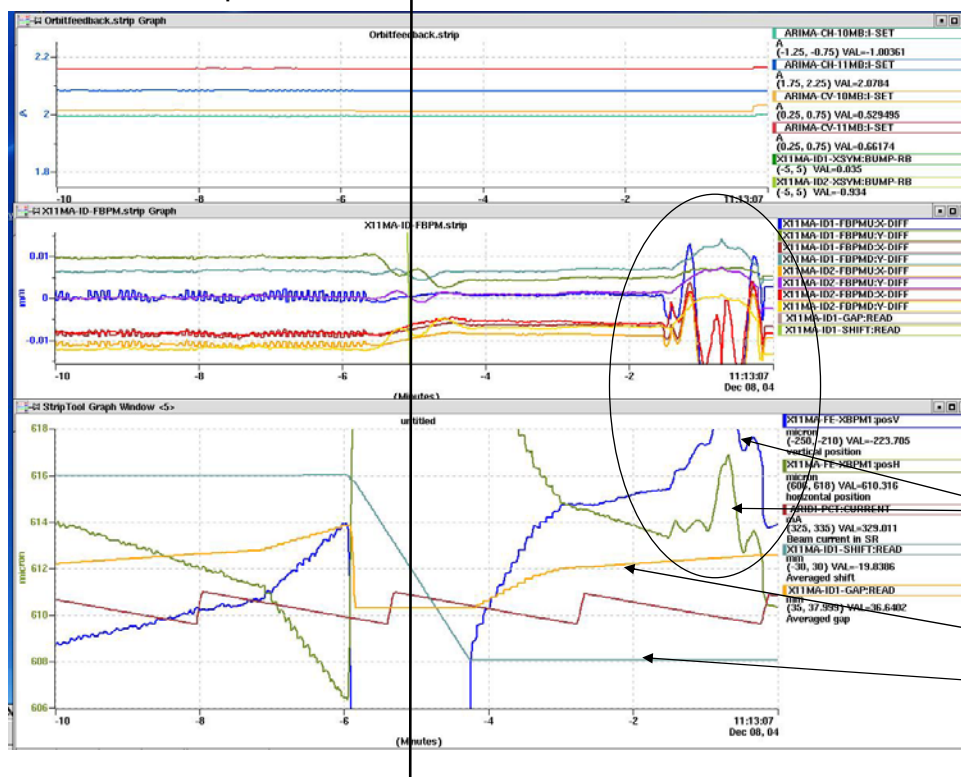


Increased noise!!!

Slow Orbit feedback

Circular plus

Circular minus

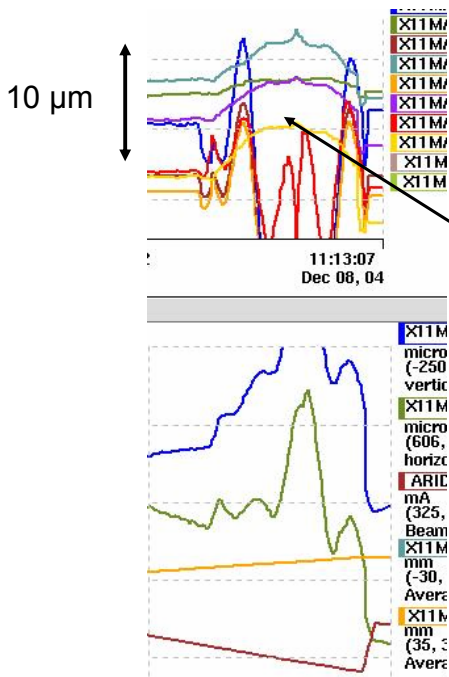


Orbit correctors?

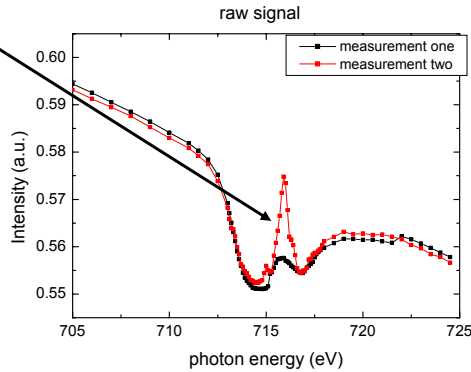
beam position in ID (Bergoz)

X-ray position in beamline

Shift and gap

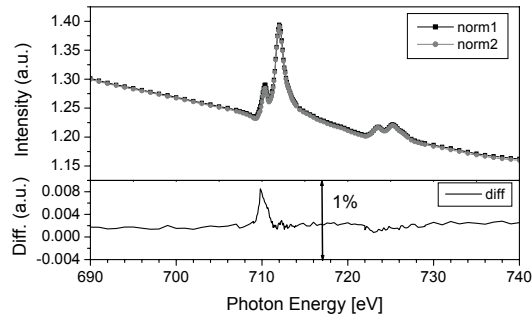


Not transparent !



Summary

We can do great measurements at SLS



- Transparent IDs are essential! Very difficult to make a double EPU system 100% transparent. Have to rely on Orbit feedback
- For the measurement “no” difference between slow and no Orbit feedback
- Critical time scale second – hour
(10 Hz – 0.0001 Hz)
Intensity variation 0.1% \approx 0.5 μ m
energy variation 1meV \approx 1 μ m
- Slow Orbit feedback is not sufficient

Fast Orbit feedback is great